

Model 2022

Spectroscopy

Amplifier

9231208C

User's Manual



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The information in this document describes the product as accurately as possible, but is subject to change without notice.

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Important Safety Considerations

Read Carefully



Indicates warning of mains or high voltage present at output labeled HV.
Risk of electrical shock if covers are removed.



Caution - risk of danger. Refer to documentation for detailed explanation of caution symbol wherever marked.



Product complies with appropriate current EU directives.



Example of "Cue" mark. Product complies with appropriate current FCC (UL)/CSA 61010-1 directives.

Manufacturer's Address

Canberra Industries, Inc.
800 Research Parkway
Meriden, CT 06450

1. Introduction

The Model 2022 Spectroscopy Amplifier provides excellent pulse symmetry, minimum sensitivity of output amplitude to variations in detector rise time, and maximum signal to noise ratio. Unipolar shaping is achieved with one differentiator and two active filter integrators. The differentiator is placed early in the amplifier to insure good overload recovery. The integrators are placed late to minimize noise contribution from the gain stages. The amplifier offers six front panel selectable pulse shaping time constants: 0.5, 1, 2, 4, 8 and 12 μ s.

The Model 2022 employs Canberra's unique baseline restorer for optimum performance with high resolution detector systems. The gated baseline restorer automatically adjusts the restoration rate and threshold optimizing performance to the incoming count rate and system noise level.

Simultaneous unipolar and bipolar outputs are available at both front and rear panel BNC connectors. The bipolar output can be used for counting, timing, or gating.

The Model 2022 borrows the Model 2020's dc stability and low noise to provide a high performance spectroscopy amplifier in a single width NIM module.

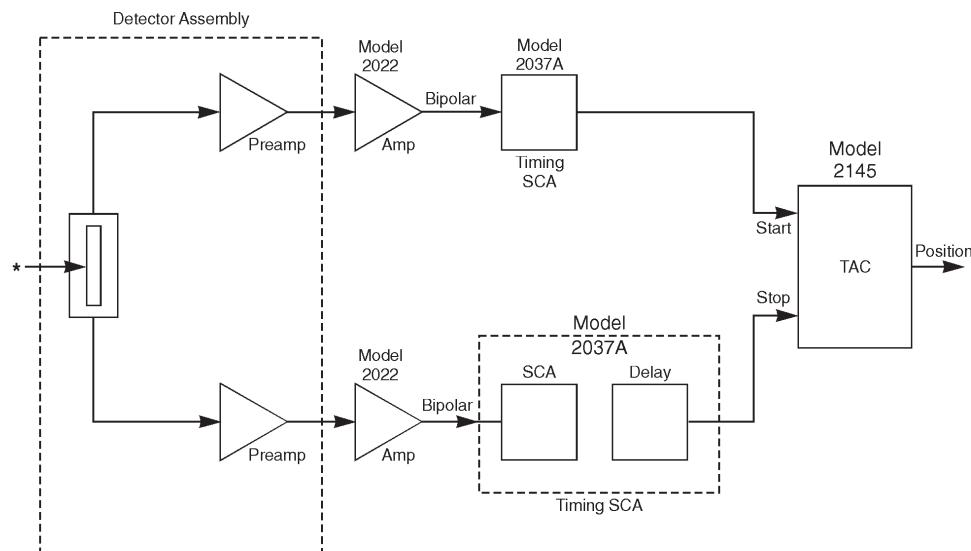


Figure 1 Position Sensitive Detector System

2. Controls and Connectors

Front Panel

This is a brief description of the 2022's front panel controls, indicators, and connectors. For more detailed information, refer to Appendix A, Specifications.

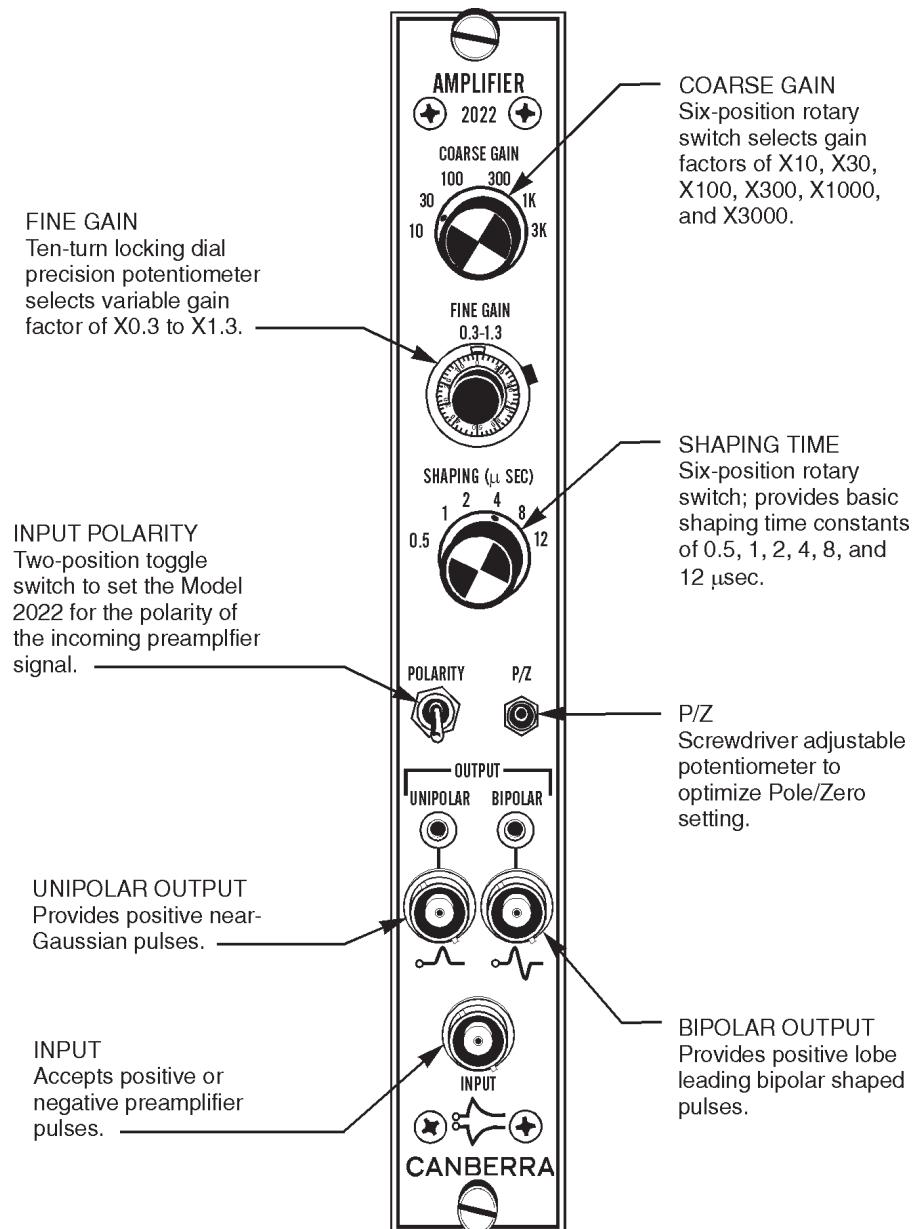


Figure 2 Front Panel Controls and Connectors

Rear Panel

This is a brief description of the 2022's rear panel connectors. For more detailed information, refer to Appendix A, Specifications.

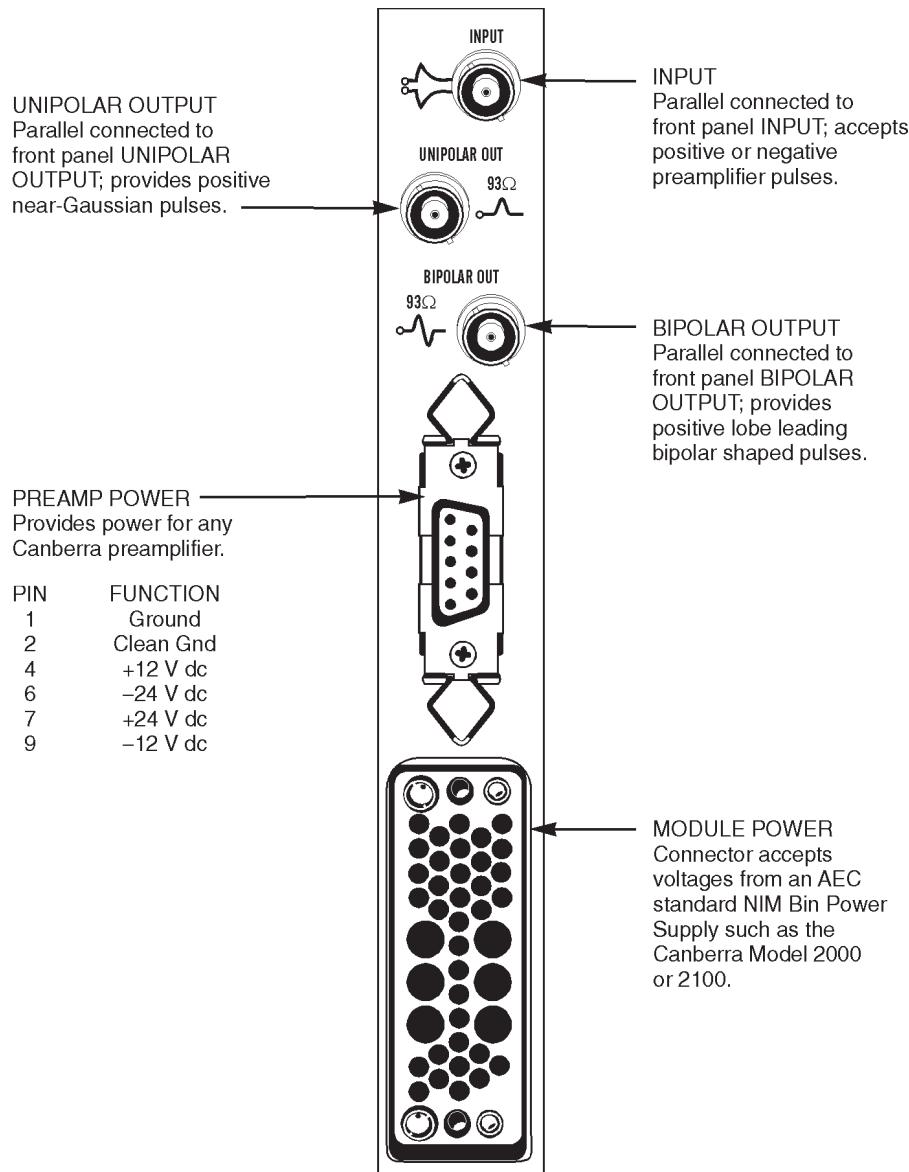


Figure 3 Rear Panel Connectors

Internal Controls

For proper operation of the Model 2022, jumpers J1 and J2 must be placed in the OUT position.

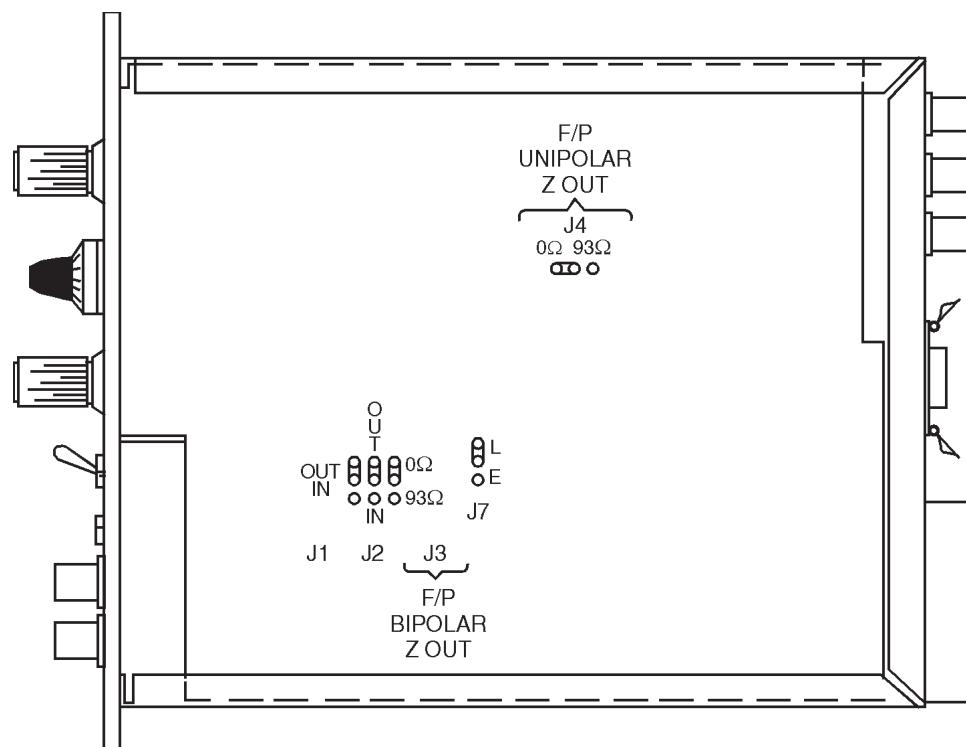


Figure 4 Internal Controls (Right Side Cover Removed)

3. Operations

The purpose of this section is to familiarize the user with the operation of the Model 2022 Amplifier and to check that the unit is functioning correctly. If the following procedures are carried out, the user will gain sufficient familiarity with this instrument to permit its proper use in the system at hand.

Installation

The Canberra Model 2000 bin and power supply, or other bin and power supply systems conforming to DOE/ER-0457T will accommodate the Model 2022. The right side cover of the two-width NIM module acts as a guide for insertion of the instrument. The module is secured in place by turning the two front-panel captive screws clockwise until finger tight. It is recommended that the NIM bin power switch be OFF whenever the module is installed or removed.

The Model 2022 can be operated where the ambient air temperature is between 0 °C and +50 °C (+120 °F maximum). Perforations in the top and bottom sides permit cooling air to circulate through the module. When relay rack mounted along with other heat generating equipment, adequate clearance should be provided to allow sufficient air flow through both the perforated top and bottom covers of the NIM bin.

Spectroscopy System Operation

A block diagram of a typical Canberra gamma spectroscopy system is shown in Figure 5.

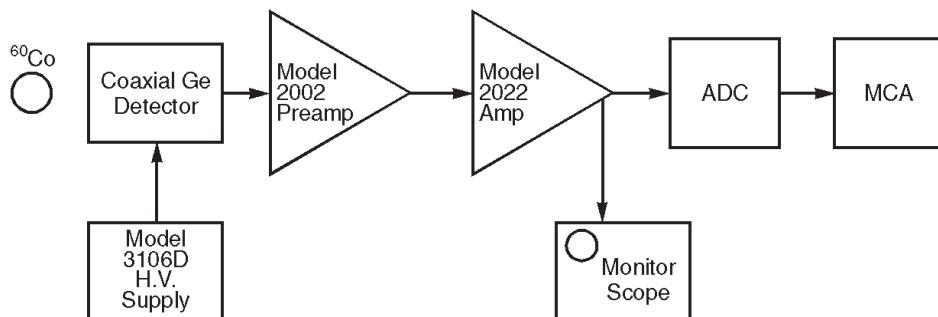


Figure 5 Typical Gamma Spectroscopy System

System Setup

Internal Controls

1. Prior to installation, the internal controls should be set to their desired positions. See Figure 4 for the locations of the controls.

There are two jumper plugs (J1 and J2) that set the UNIPOLAR output to the prompt (OUT) mode, these jumper should be left in the OUT position.

There are two jumper plugs which select the output impedance of the front panel (only) UNIPOLAR (J4) and BIPOLAR (J3) outputs. The output impedances can be changed from 0 ohms to approximately 93 ohms. The Model 2022 is shipped with the output impedance set for 0 ohms. The rear panel outputs have a fixed impedance of 93 ohms, series connected.

When using the front panel low impedance output, short lengths of interconnecting coaxial cable need not be terminated. To prevent possible oscillations, longer cable lengths should be terminated at the receiving end in a resistive load equal to the cable impedance (93 ohms for type RG-62 cable).

The 93 ohm output may be safely used with RG-62 cable up to a few hundred meters. However, the 93 ohm impedance is in series with the load impedance, and a decrease in the total signal range may occur. For example, a 50% loss will result if the load impedance is 93 ohms.

Another jumper plug (J7) is provided to select a linear or an exponential restorer response. The linear response is optimum for the faster shaping-time constants and their associated noise spectra. The slower shaping-time constants produce noise spectra having proportionately lower frequency components, requiring a different restorer response. The exponential response is better suited to these shapings.

The Model 2022 is shipped with the restorer response jumper plug in the linear (L) response position, which will give the best restorer response for shaping-time constants of 0.5, 1, 2, and 4 μ s. These are the time constants generally used with Ge, proportional counter, surface barrier, and scintillation photomultiplier detectors.

The exponential (E) response position yields optimum restorer performance with shaping-time constants of 8 and 12 μ s. These time constants are most often used with Planar Ge and Si(Li) detectors. The exponential response may also be used with the other shaping time constants without loss of resolution performance, provided that the input count-rate remains below 20 kcps.

2. Insert the Model 2022 into a standard NIM bin. Preamp power is provided by means of a connector located on the rear panel of the Model 2022 amplifier. Allow the entire system to warm up and stabilize.

3. Set the Model 2022 controls as follows:

Shaping:	2 μ s
Coarse Gain:	100
Fine Gain:	7.2

4. Set the INPUT POLARITY switch to match the output polarity of the preamp (positive, +, for a Canberra Model 2001 Preamp). This will give approximately a 9V output when using a preamp gain of 100 mV/MeV and a ^{60}Co source.
5. Install a “Tee” connector on the Model 2022 UNI POLAR output. Connect one end to the ADC’s INPUT. The ADC must be direct coupled for linear input signals to fully exploit the rate capabilities of the Model 2022. All Canberra ADCs are dc coupled.
6. Connect the other end of the “Tee” connector to an oscilloscope to monitor the UNIPOLAR output.

Performance Adjustments

The pole/zero trim, which is extremely critical for good high count-rate resolution performance, can be adjusted in either of two ways. The first used a detector and a standard source or radiation. The other used a square wave generator. Each of these is described in the following paragraphs.

Pole/Zero Using a Ge Detector and ^{60}Co

Adjust the source’s count to between 2 kcps and 25 kcps. While observing UNIPOLAR output on the scope, adjust the pole/zero so that the trailing edge of the unipolar pulse returns to the baseline with no over- or undershoots. Figure 6 shows the correct setting of the P/Z control, with Figures 7 and 8 showing under- and over-compensation for the preamplifier decay time-constant. Notice some small amplitude signals with long decay times in Figure 6. These are due to charge trapping in the detector and cannot be corrected by the P/Z control.

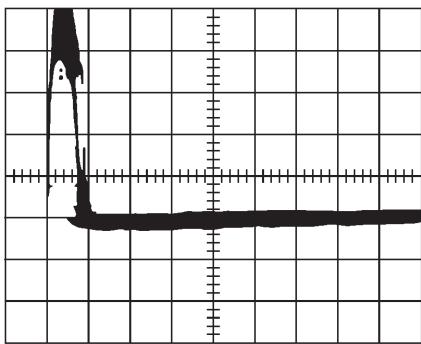


Figure 6 Correct Pole/Zero Compensation

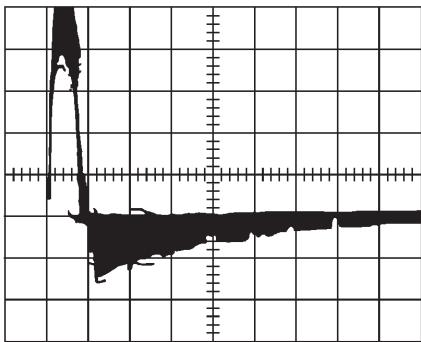


Figure 7 Undercompensated Pole/Zero

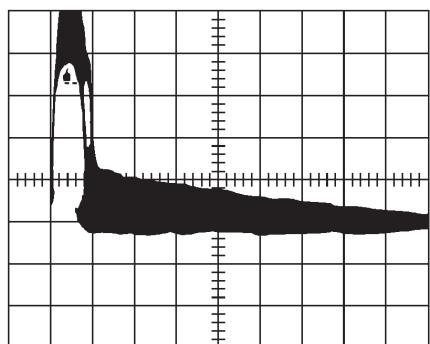


Figure 8 Overcompensated Pole/Zero

Oscilloscope
Vert: 50 mV/div
Horiz: 10 μ s/div

Source ^{60}Co
1.33 MeV peak: 7 V amplifier
Count Rate: ≈ 2 kcps
Shaping: 2 μ s

Pole/Zero Adjustment Using a Square Wave and Preamplifier Test Input

1. Driving the preamp test input with a square will allow a more precise adjustment of the amplifier P/Z.
2. The amplifier's controls should basically be set for its intended application: COARSE GAIN, SHAPING, INPUT POLARITY.
3. Adjust the square wave generator for a frequency of approximately 1 kHz.
4. Connect the square wave generator's output to Preamp's TEST INPUT.
5. Remove all radioactive sources from the vicinity of the detector
6. Set the scope's channel 1 vertical sensitivity to 5 V/div and adjust the main time base to 0.2 ms/div. Monitor the Model 2022 UNIPOLAR output and adjust the square wave generator's amplitude control (attenuator) for output signals of ± 8 V.

Note: Both positive and negative unipolar pulses will be observed at the output.

7. Reduce the scope vertical sensitivity to 50 mV/div. See the note below.

Note: At high count rates, the Pole/Zero adjustment is extremely critical for maintaining good resolution and low peak-shift. For a precise and optimum setting for the Pole/Zero, a scope vertical sensitivity of 50 mV/div should be used. Higher scope sensitivities can also be used, but result in a less precise Pole/Zero adjustment. However, most scopes will over load for a 10 V input signal when vertical sensitivity is set for 50 mV/div. Overloading the scope input will distort the signals' recovery to the baseline. Thus the Pole/Zero will be incorrectly adjusted, thereby resulting in a loss of resolution at high count-rates. To prevent overloading the scope, a clamping circuit, such as Canberra Model LB1502 Schottky Clamp Box, should be used.

Figure 9 shows the correct setting of the Pole Zero control. Figures 10 and 11 show under- and over-compensation for the preamplifier decay-time constant. As illustrated in Figure 9, the UNIPOLAR output signal should have a clean return to the baseline with no bumps, overshoots or undershoots.

Note: When adjusting the P/Z using the square wave technique, the calibration square wave generated by the oscilloscope can be used. Most scopes generate a 1 kHz square wave used to calibrate vertical gain and probe compensation. Connect the scope CALIBRATION Output through an attenuator to the preamp input and repeat steps one through six in "Performance Adjustments" on page 7.

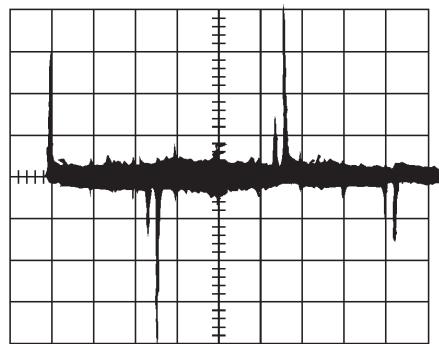


Figure 9 Correct Pole/Zero Compensation

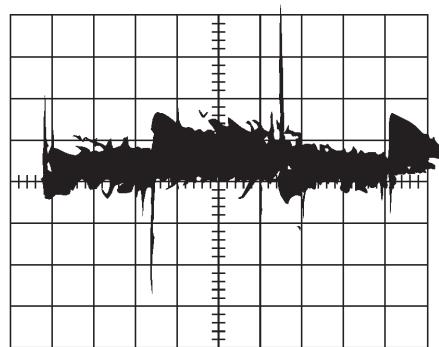


Figure 10 Undercompensated Pole/Zero

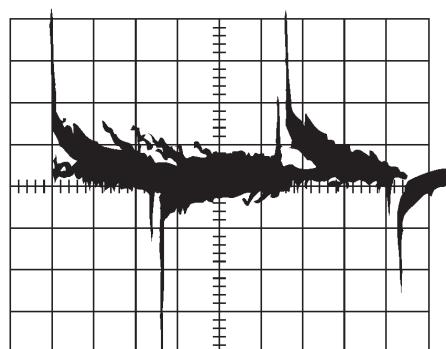


Figure 11 Overcompensated Pole/Zero

Restorer Response

Faster shaping-time constants (0.5 through 4 μ s) give rise to output noise spectra having different frequency components than the slower shaping-time constants (5 through 12 μ s). The Model 2022 provides a restorer response for each of the two conditions: linear (L) response for the faster time constants and an exponential (E) response for the slower time constants. Jumper plug J7 (see figure 4) selects the required response. Shipped in the linear (L) position. See “System Setup, Internal Controls” on page 6 for more information.

MCA Controls

To obtain optimum resolution, the Lower Level Discriminator (LLD) on the MCA/ADC should be set just above the noise so that the effects of pileup are minimized.

Resolution Versus Count Rate and Shaping

The 2 μ s shaping-time constant is near-optimum for Ge detector systems over a wide range of incoming count rates. For high resolution, larger shaping-time constants offer a better signal-to-noise (S/N) ratio, resulting in better resolution. However, as the count rate increases, the effects of pileup will degrade the resolution much sooner. The optimum shaping-time constant depends on the detector characteristics (such as its size, configuration, and collection characteristics), preamplifier, and incoming count rate.

Below is a list of the optimum shaping-time constant ranges for other common detectors:

Scintillation Photomultiplier	0.5
Gas Proportional Counters	0.5 through 2
Silicon Charged Particle	0.5 through 2
Coaxial Germanium	2 through 4
Planar Germanium	4 through 12
Cooled Silicon.	8 through 12

Resolution Destroying Interferences

Vibration Transmitted to Detector and Cryostat

This can be through the floor or mounting, as well as direct audio coupling through the air. Vibration isolators in the mounting and sound absorbing covers around the detector can reduce this problem.

Close Proximity of a Radio Station Picked Up by the Cryostat’s “Dipstick”

Good contact between the dipstick and the cryostat can often help solve this problem. Beware of grounding the cryostat and dipstick because this may increase power line frequency (50 or 60 Hz) ground loops.

Ground Loops and Power Line Frequency Interference Caused by Long Cable Connections Between Detector, Preamplifier and Shaping Amplifier

There is no general solution for this problem. As a first step, the preamp should use the power supplied by the main shaping amplifier. Second, the system should have a single point house ground. For example, on a general system connect the NIM bin to house ground via the ac line cord. Isolate all other equipment requiring ac voltage from the house ground. Connect all the chassis in the system to the grounded NIM bin using heavy braided wire.

High Voltage Power Supplies (HVPS)

Generally, an ac line powered HVPS should float from power line ground with the only ground being made at the preamplifier through the high voltage connecting cable.

Analyzer EMI

If the detector is located within three to five meters (10 to 15 feet) of a multichannel analyzer containing ferrite CORE memory, it can receive EMI (electromagnetic interference). This is due to high memory-core-currents during the memory cycle of the analyzer. The only practical cure for this problem is to operate the analyzer in the “Live” Mode of accumulation. That way, the memory cycle operates only while no signal is being analyzed.

Amplifier Parasitic Oscillations

If the cable connecting the front panel outputs of the amplifier to the ADC exceed three to six meters (10 to 15 feet) in length, oscillations can occur. The cure is to use RG-62 cable (93 ohm impedance) and terminate the ADC end of the cable with a 93 ohm metal film resistor. Alternatively, the 93 ohm output impedance of the amplifier can be used with no terminator.

Trouble Isolated to Ground Loops and or RF EMI

If trouble is isolated to ground loops and or RF EMI, detector-system loop-buster accessories are available to help minimize these effects. An application note entitled “System Considerations with High Resolution Detectors” is available from the factory.

A. Specifications

Inputs

INPUT – Accepts positive or negative pulses from an associated preamplifier; amplitude: ± 10 V divided by the selected gain for linear response; ± 12 V maximum; rise time: less than SHAPING time constant; decay time constant; $40 \mu\text{s}$ to ∞ for 0.5, 1, 2, 4 and $8 \mu\text{s}$ shaping time constants, $100 \mu\text{s}$ to ∞ for $12 \mu\text{s}$ shaping time constant; $Z_{\text{in}} \approx 1 \text{k}\Omega$; front and rear panel BNC connectors.

Outputs

UNIPOLAR OUTPUT – Provides positive linear active-filtered near-Gaussian shaped pulses; amplitude linear to $+10$ V, 12 V max.; dc restored; output dc level factory calibrated to 0 ± 5 mV, front panel $Z_{\text{out}} < 1 \Omega$ or $\approx 93 \Omega$, internally selectable; rear panel $Z_{\text{out}} \approx 93 \Omega$; short circuit protected; front and rear panel BNC connectors.

BIPOLAR OUTPUT – Provides prompt positive lobe leading linear active-filtered bipolar-shaped pulses; amplitude linear to $+10$ V, 12 V max., negative lobe is approximately 70% of positive lobe; dc coupled; output dc level ± 25 mV; front panel $Z_{\text{out}} < 1 \Omega$ or $\approx 93 \Omega$, internally selectable; rear panel $Z_{\text{out}} \approx 93 \Omega$; short circuit protected; front and rear panel BNC connectors.

Front Panel Controls

COARSE GAIN – Rotary switch selects gain factors of X10, X30, X100, X300, X1000 and X3000.

FINE GAIN – Ten-turn locking-dial precision potentiometer selects variable gain factor of X0.3 to X1.3; resetability $\leq 0.03\%$.

INPUT POLARITY – Toggle switch selects the polarity of the incoming preamplifier signal.

P/Z – Multi-turn screwdriver adjustable pole/zero potentiometer optimizes amplifier baseline recovery and overload performance for the preamplifier fall time constant and the 2022's pulse shaping chosen; $40 \mu\text{s}$ to ∞ for 0.5, 1, 2, 4 and $8 \mu\text{s}$ SHAPING time constants, $100 \mu\text{s}$ to ∞ for $12 \mu\text{s}$ SHAPING time constant.

SHAPING TIME – Rotary switch provides 0.5, 1, 2, 4, 8 and 12 μs basic shaping time constants.

Internal Controls

UNIPOLAR Z_{out} – Jumper plug provides $Z_{\text{out}} \leq 1 \Omega$ or $\approx 93 \Omega$ for the front panel UNIPOLAR output. Shipped in the $\leq 1 \Omega$ position.

BIPOLAR Z_{out} – Jumper plug provides $Z_{\text{out}} \leq 1 \Omega$ or $\approx 93 \Omega$ for the front panel BIPOLAR output. Shipped in the $\leq 1 \Omega$ position.

L/E – Jumper plug selects a linear or exponential restorer response. Shipped in the L (linear) position.

Performance

GAIN RANGE – Continuously variable from X3 to X3900, product of COARSE and FINE GAIN controls.

GAIN DRIFT – $\leq \pm 0.0075\%/\text{^{\circ}C}$.

DC LEVEL DRIFT – UNIPOLAR output: $\leq \pm 10 \mu\text{V}/\text{^{\circ}C}$; BIPOLAR OUTPUT: $\leq \pm 50 \mu\text{V}/\text{^{\circ}C}$.

INTEGRAL NON-LINEARITY – $\leq \pm 0.05\%$, over total output range for 2 μs shaping.

CROSSOVER WALK – BIPOLAR output: $\leq \pm 3 \text{ ns}$ for 50:1 dynamic range and 2 μs shaping when used with Canberra Model 2037A Edge/Crossover Timing Single Channel Analyzer.

OVERLOAD RECOVERY – UNIPOLAR (BIPOLAR) output recovery to within $\pm 2\%$ (1%) of full scale output from X1000 overload in 2.5 (2.0) non-overloaded pulse widths, at full gain, any shaping time constant and pole/zero cancellation properly set.

NOISE CONTRIBUTION – $\leq 4.0 \mu\text{V}$ true RMS UNIPOLAR ($7.1 \mu\text{V}$ BIPOLAR) output referred to input, 2 μs shaping and amplifier gain ≥ 100 .

PULSE SHAPING – Near-Gaussian shape; one differentiator (two for bipolar), two active filter integrators; UNIPOLAR time to peak: 2.35X shaping time; pulse width: 7.3X shaping time BIPOLE time to crossover: 2.8X shaping time, time to peak, pulse width and crossover times measured at 0.1% of full scale output; 1 μ s SHAPING center frequency: 150 kHz; band width: 180 kHz; f_c and BW for other shaping are multiples of that given for 1 μ s.

RESTORER – Active gated.

SPECTRUM BROADENING – The FWHM of a ^{60}Co 1.33 MeV gamma peak for an incoming rate of 2 kcps to 100 kcps and a 9 V pulse height will typically change <14% for 2 μ s shaping. These results may not be reproducible if associated detector exhibits an inordinate amount of long rise time signals.

COUNT RATE STABILITY – The peak position of a ^{60}Co 1.33 MeV gamma peak for an incoming count rate of 2 kcps to 100 kcps and a 9 V pulse height will typically shift <0.024% for 2 μ s shaping.

Connectors

All signal connectors are BNC type.

PREAMP POWER – Rear panel, Amphenol type 17-10070.

Power Requirements

+24 V dc – 125 mA +12 V dc – 75 mA

-24 V dc – 150 mA -12 V dc – 65 mA

Physical

SIZE – Standard single-width NIM module 3.43 x 22.12 cm (1.35 x 8.71 in.) per DOE/ER-0457T.

NET WEIGHT – 0.9 kg (2.0 lb).

SHIPPING WEIGHT – 1.9 kg (4.1 lb).

Environmental

OPERATING TEMPERATURE RANGE - 0 to 50 °C (32 to 122 °F).

OPERATING HUMIDITY – 0 to 80% relative, non-condensing.

Meets the environmental conditions specified by EN 61010, Installation Category I,
Pollution Degree 2.

B. Environmental Considerations

This unit complies with all applicable European Union requirements.

Compliance testing was performed with application configurations commonly used for this module; i.e. a CE compliant NIM Bin and Power Supply with additional CE compliant application-specific NIM were racked in a floor cabinet to support the module under test.

During the design and assembly of the module, reasonable precautions were taken by the manufacturer to minimize the effects of RFI and EMC on the system. However, care should be taken to maintain full compliance. These considerations include:

- A rack or tabletop enclosure fully closed on all sides with rear door access
- Single point external cable access
- Blank panels to cover open front panel Bin area
- Compliant grounding and safety precautions for any internal power distribution
- The use of CE compliant accessories such as fans, UPS, etc.

Preventive Maintenance

This unit does not require any periodic cleaning maintenance.

Any maintenance should be performed by a qualified Canberra service representative.

Operating Protection Impairment

Canberra is not liable for any operational malfunctions or personal injuries due to mis-handling or unauthorized repair and maintenance not detailed in this manual.



Cleaning/Decontamination

When needed, the unit may be cleaned. Remove power from the unit before cleaning. Use only a soft cloth dampened with warm water and do not allow water to enter the unit. Make sure unit is fully dry before restoring power. Because of excess holes in the NIM wrap, *do not* use any liquids to clean the wrap, side, or rear panels.

Notes

Request for Circuit Information

The Schematics, Block Diagrams and/or Circuit Description may be available for this unit directly from CANBERRA. Request can be made by calling, faxing, or emailing:

Training and Technical Services Department
Canberra Industries
800 Research Parkway, Meriden, CT 06450
Telephone: (800) 255-6370
FAX: (203) 639-2067
Email: techsupport@canberra.com

If you would like schematics and/or a circuit description, if available, for this unit, please provide us with the following information.

Your Name _____

Your Address _____

Unit's model number _____

Unit's serial number _____

Note: Schematics, block diagrams, and circuit descriptions are provided for information only; if you service or repair or try to service or repair this unit without Canberra's written permission you may void your warranty.

Notes

Warranty

Canberra (we, us, our) warrants to the customer (you, your) that for a period of ninety (90) days from the date of shipment, software provided by us in connection with equipment manufactured by us shall operate in accordance with applicable specifications when used with equipment manufactured by us and that the media on which the software is provided shall be free from defects. We also warrant that (A) equipment manufactured by us shall be free from defects in materials and workmanship for a period of one (1) year from the date of shipment of such equipment, and (B) services performed by us in connection with such equipment, such as site supervision and installation services relating to the equipment, shall be free from defects for a period of one (1) year from the date of performance of such services.

If defects in materials or workmanship are discovered within the applicable warranty period as set forth above, we shall, at our option and cost, (A) in the case of defective software or equipment, either repair or replace the software or equipment, or (B) in the case of defective services, reperform such services.

LIMITATIONS

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EXCLUSIONS

Our warranty does not cover damage to equipment which has been altered or modified without our written permission or damage which has been caused by abuse, misuse, accident, neglect or unusual physical or electrical stress, as determined by our Service Personnel.

We are under no obligation to provide warranty service if adjustment or repair is required because of damage caused by other than ordinary use or if the equipment is serviced or repaired, or if an attempt is made to service or repair the equipment, by other than our Service Personnel without our prior approval.

Our warranty does not cover detector damage due to neutrons or heavy charged particles. Failure of beryllium, carbon composite, or polymer windows, or of windowless detectors caused by physical or chemical damage from the environment is not covered by warranty.

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